

Gaussian Mixture Models for Inversion and Uncertainty Quantification

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Gaussian Mixture Inversion

Application to OCO-2



Introduction

Terrestrial Carbon Cycle



Figure 1: NASA/JPL



Carbon Flux Mismatch



Figure 2: globalcarbonproject.org/carbonbudget/index.htm



Carbon Flux Inversion



Figure 3: NASA/JPL



The Orbiting Carbon Observatory 2 (OCO-2) Instrument



Figure 4: NASA/JPL



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The Orbiting Carbon Observatory 2 (OCO-2) Instrument

Measured radiances are used to infer atmospheric CO₂ concentrations.









The OCO-2 mission uses Optimal Estimation (Rodgers, 2000). The relationship between atmospheric state vector *x* and radiance vector *y* is modeled as

$$y = F(x, b) + \varepsilon, \quad x \sim N(x_a, S_a), \quad \varepsilon \sim N(0, S_y).$$
 (1)

The atmospheric state is then inferred by solving the corresponding inverse problem:

$$\widehat{x} = \operatorname*{argmin}_{x} \left[y - F(x,b) \right]^{T} S_{y}^{-1} \left[y - F(x,b) \right] + \left[x - x_{a} \right]^{T} S_{a}^{-1} \left[x - x_{a} \right].$$
(2)

The operational retrieval also provides an estimate for the posterior covariance:

$$\widehat{\mathsf{S}} = \left(\mathsf{K}^{\mathsf{T}}\mathsf{S}_{\mathsf{y}}^{-1}\mathsf{K} + \mathsf{S}_{a}^{-1}\right)^{-1}.$$
(3)



Ground Measurement: TCCON





- -Retrievals take significant computational effort. (Re-)Processing the entire OCO-2 data record takes more than a year.
- -OE doesn't provide trustworthy uncertainty estimates. Posterior distribution might be non-Gaussian, overall estimate is generally too low. Error sources outside retrieval algorithm are not included.
- -XCO2 estimates are biased. Bias correction needed after inital processing.



Towards Solutions

-Direct retrievals using neural networks (David et al. [2021], Bréon et al [2022]) offer a fast means to going to XCO2 directly from input radiances.



Figure 5: Image from David et al.

-MCMC aided by surrogate forward model. (Lamminpää et al. [2019] for exploring the non-Gaussian posterior. -Simulation Based UQ (Braverman et al. [2021]) simultaneously tackles bias and non-Gaussianity.



Figure 6: Left: Lamminpää et al., Right: Braverman et al.



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Gaussian Mixture Inversion

Modeling a Non-Linear Function u = f(w)



Figure 7: Image from Braverman et al.



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Gaussian Mixture Model: Joint Distribution

$$p(y,x) = p(z) = \sum_{k=1}^{K} \pi^{(k)} \phi(z; \mu^{(k)}, \Sigma^{(k)})$$

$$z = \begin{bmatrix} y \\ x \end{bmatrix}, \quad \Sigma^{(k)} = \begin{bmatrix} \Sigma_{yy}^{(k)} & \Sigma_{yx}^{(k)} \\ \Sigma_{xy}^{(k)} & \Sigma_{xx}^{(k)} \end{bmatrix}$$
(5)





Gaussian Mixture Model: Conditional Distribution

$$p(x|y) = \sum_{k=1}^{K} \pi_{x|y}^{(k)} \phi(x; \mu_{x|y}^{(k)}, \Sigma_{x|y}^{(k)})$$
(6)

$$\pi_{x|y}^{(k)} = \frac{\pi^{(k)}\phi(y;\mu_y^{(k)},\Sigma_{yy}^{(k)})}{\sum_{l=1}^{K}\pi^{(l)}\phi(y;\mu_y^{(l)},\Sigma_{yy}^{(l)})}$$
(7)

$$\mu_{X|y}^{(k)} = \mu_{X}^{(k)} + \Sigma_{Xy}^{(k)} \left(\Sigma_{yy}^{(k)}\right)^{-1} \left[y - \mu_{y}^{(k)}\right]$$
(8)

$$\Sigma_{x|y}^{(k)} = \Sigma_{xx}^{(k)} - \Sigma_{xy}^{(k)} \left(\Sigma_{yy}^{(k)}\right)^{-1} \Sigma_{yx}^{(k)}$$
(9)



A Bayesian solution to an non-linear inverse problem of solving for x in $y = F(x) + \varepsilon$ is given by

$$p(x|y) \propto p(y|x)p(x).$$
 (10)

Instead of gradient-based solutions, we propose *Gaussian Mixture Inversion* (GMI): given $x \sim p(x)$ and $y = F(x) + \varepsilon$, the posterior density will be approximated as

$$\widetilde{p}(x|y) = \sum_{k=1}^{K} \pi_{x|y}^{(k)} \phi(x; \mu_{x|y}^{(k)}, \Sigma_{x|y}^{(k)})$$
(11)

defined as before.



Benchmark solution: Adaptive Metropolis (AM). Sample points x_{t+1} from a proposal distribution $N(x_t, C_t)$, where C_t is the covariance matrix of the chain at time t:

$$C_t = cov([x_1, \dots, x_t]), \tag{12}$$

accept new point with probability

$$\alpha(\mathbf{x}_t, \mathbf{x}_{t+1}) = \min\left(1, \frac{\pi(\mathbf{x}_{t+1})}{\pi(\mathbf{x}_t)}\right)$$
(13)

Further, we observe

$$\mathbb{E}^{\mathsf{Y}}\left[d_{Hell}\left(\pi(\cdot|\mathsf{Y}),\pi_{\mathsf{N}}(\cdot|\mathsf{Y})\right)^{2}\right] \leq 2\left(d_{Hell}(\pi(\cdot,\cdot),\pi(\cdot,\cdot)_{\mathsf{N}})^{2} + d_{Hell}(\pi_{\mathsf{Y}},\pi_{\mathsf{Y}_{\mathsf{N}}})^{2}\right)$$
(14)



We validate our approach to posterior approximation by considering the following two synthetic example. Let

$$x \in \mathbb{R}, \quad y = f(x) + \varepsilon \in \mathbb{R}, \quad f(x) = \sin(2x) - \cos(3x)$$
 (15)



Figure 8: Test case where true x = -0.4



2D Toy Example

Next, consider

$$x \in \mathbb{R}^2$$
, $y = f(x) + \varepsilon \in \mathbb{R}$, $f(x) = \sin(x_1) - \cos(x_2)$ (16)



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Application to OCO-2

Example Using Simulated OCO-2 Data

- Retrieval for a single pixel. Quantity of interest: column averaged CO2 concentration, denoted XCO2.
- State vector x with realistic atmospheric and surface conditions.
- Simulated measurement: forward model evaluated at *x*, add synthetic measurement error.
- Radiance dimension reduction using PCA.
- Compare GMI and MCMC posteriors (with and without "model discrepancy", using a forward model emulator. See talk by Jouni Susiluoto, MS157, on Thursday!).



$$y = F(x, b) + \varepsilon + \delta \tag{17}$$

- *x*: operational prior, combination of priors, more comprehensive local / global distribution.
- *b*: forward model parameter uncertainty included by prescribing a distribution.
- $\cdot \ \varepsilon:$ error model, possibility of off-diagonal elements in covariance, can be non-Gaussian.
- δ: model discrepancy for including model misspecification, can include new ML bias-correction, other methods for accounting "Unknown Unknowns" and spectral residuals.



To test performance with more realistic data, we perturb modeled radiances with model discrepancy adjustment:

$$y = F_1(x, b_1) + \varepsilon + \delta, \quad \delta = F_0(x, b_0) - F_1(x, b_1).$$
 (18)



Left: O2-A Band radiance. Middle: a realization of measurement error $\varepsilon \sim \mathcal{N}(0, S_y)$. Right: a realization of model discrepancy $\delta \sim \mathcal{N}(\mu_{\delta}, S_{\delta})$.



Example Using Simulated OCO-2 Data



(Without model discrepancy: $y_{obs} = F(x) + \varepsilon$)



Full posterior



Example Using Simulated OCO-2 Data



(With Model Discrepancy: $y_{obs} = F(x) + \varepsilon + \delta$)



Full posterior, with model discrepancy



Timing

- 10000 state vectors sampled from prior; input to forward model (emulator); added noise.
 Time: 104s (parallel in DGX station with 20 threads).
- Traned GMM with Julia's GaussianMixtures package: 100 iterations for learning, 20 mixture components. Time: 44s.
- Condition on measured radiance; sample 10000 realizations from conditional distribution. Time: 0.014s.
- Adaptive Metropolis ran for 1000000 iterations. Time: approx. 17000s, or 4.5h (using emulator).



- Training data: partition the Globe into clusters according to real OCO-2 measurements via self-organizing maps. For each cluster, derive marginal distribution on x and δ. Obtain y = F(x, b) + ε + δ.
- Evaluate model performance against operational retrieval by left-out simulated data, and TCCON co-located real world retrievals of XCO2.



Comparison with Left-Out Data





Comparison with Real Data





Include Model Discrepancy and Prior Mean In Training





Comparison with Real Data (Revisited)





Remark: Averaging Kernels?

$$\begin{aligned} \mathbf{A} &= (\mathbf{S}_{a}^{-1} + \mathbf{K}^{\mathsf{T}} \mathbf{S}_{\varepsilon}^{-1} \mathbf{K})^{-1} \mathbf{K}^{\mathsf{T}} \mathbf{S}_{\varepsilon}^{-1} \mathbf{K} \\ &= \widehat{\mathbf{S}} \mathbf{K}^{\mathsf{T}} \mathbf{S}_{\varepsilon}^{-1} \mathbf{K} \\ &= \widehat{\mathbf{S}} (\mathbf{K}^{\mathsf{T}} \mathbf{S}_{\varepsilon}^{-1} \mathbf{K} + \mathbf{S}_{a}^{-1} - \mathbf{S}_{a}^{-1}) \\ &= \widehat{\mathbf{S}} (\widehat{\mathbf{S}}^{-1} - \mathbf{S}_{a}^{-1}) \\ &= \widehat{\mathbf{S}} \widehat{\mathbf{S}}^{-1} - \widehat{\mathbf{S}} \mathbf{S}_{a}^{-1} \\ &= \mathbf{I} - \widehat{\mathbf{S}} \mathbf{S}_{a}^{-1} \end{aligned}$$
(19)



- Wasserstein loss and cross-validation to fit GMM.
- Fitting/learning the mixture model with e.g. mixture density networks. Different basis functions, e.g. Cauchy.
- Dimension reduction for data using Autoencoders, UMAP.
- Other methods with similar capabilities: Deep Ensemble methods, VAEs, GAN-FLOW.



Thank you!

References:

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