

Gaussian Mixture Models for Inversion and Uncertainty Quantification

Otto Lamminpää*[∗]* , Amy Braverman*[∗]* , Maggie Johnson *[∗]* , Jonathan Hobbs*[∗]* , Jouni Susiluoto*[∗]* , Mike Turmon*[∗]* February 28, 2024

* Uncertainty Quantification and Statistical Analysis Group Jet Propulsion Laboratory, California Institute of Technology [Introduction](#page-2-0)

[Gaussian Mixture Inversion](#page-13-0)

[Application to OCO-2](#page-21-0)

[Introduction](#page-2-0)

Terrestrial Carbon Cycle

Figure 1: NASA/JPL

Carbon Flux Mismatch

Figure 2: globalcarbonproject.org/carbonbudget/index.htm

Carbon Flux Inversion

Figure 3: NASA/JPL

The Orbiting Carbon Observatory 2 (OCO-2) Instrument

The Orbiting Carbon Observatory 2 (OCO-2) Instrument

Measured radiances are used to infer atmospheric $CO₂$ concentrations.

The OCO-2 mission uses Optimal Estimation (Rodgers, 2000). The relationship between atmospheric state vector *x* and radiance vector *y* is modeled as

$$
y = F(x, b) + \varepsilon, \quad x \sim N(x_a, S_a), \quad \varepsilon \sim N(0, S_y). \tag{1}
$$

The atmospheric state is then inferred by solving the corresponding inverse problem:

$$
\widehat{x} = \underset{x}{\text{argmin}} \left[y - F(x, b) \right]^T S_y^{-1} \left[y - F(x, b) \right] + \left[x - x_a \right]^T S_a^{-1} \left[x - x_a \right]. \tag{2}
$$

The operational retrieval also provides an estimate for the posterior covariance:

$$
\widehat{S} = (K^T S_y^{-1} K + S_a^{-1})^{-1}.
$$
\n(3)

Ground Measurement: TCCON

- -Retrievals take significant computational effort. (Re-)Processing the entire OCO-2 data record takes more than a year.
- -OE doesn't provide trustworthy uncertainty estimates. Posterior distribution might be non-Gaussian, overall estimate is generally too low. Error sources outside retrieval algorithm are not included.
- -XCO2 estimates are biased. Bias correction needed after inital processing.

Towards Solutions

-Direct retrievals using neural networks (David et al. [2021], Bréon et al [2022]) offer a fast means to going to XCO2 directly from input radiances.

Figure 5: Image from David et al.

-MCMC aided by surrogate forward model. (Lamminpää et al. [2019] for exploring the non-Gaussian posterior. -Simulation Based UQ (Braverman et al. [2021]) simultaneously tackles bias and non-Gaussianity.

Figure 6: Left: Lamminpää et al., Right: Braverman et al.

Jet Propulsion Laboratory 10/34 Results

[Gaussian Mixture Inversion](#page-13-0)

Modeling a Non-Linear Function $u = f(w)$

Figure 7: Image from Braverman et al. [2021]

Gaussian Mixture Model: Joint Distribution

$$
p(y, x) = p(z) = \sum_{k=1}^{K} \pi^{(k)} \phi(z; \mu^{(k)}, \Sigma^{(k)})
$$
(4)

$$
z = \begin{bmatrix} y \\ x \end{bmatrix}, \quad \Sigma^{(k)} = \begin{bmatrix} \Sigma_{yy}^{(k)} & \Sigma_{yx}^{(k)} \\ \Sigma_{xy}^{(k)} & \Sigma_{xx}^{(k)} \end{bmatrix}
$$
(5)

Gaussian Mixture Model: Conditional Distribution

$$
p(x|y) = \sum_{k=1}^{K} \pi_{x|y}^{(k)} \phi(x; \mu_{x|y}^{(k)}, \Sigma_{x|y}^{(k)})
$$
(6)

$$
\pi_{x|y}^{(k)} = \frac{\pi^{(k)}\phi(y;\mu_y^{(k)},\Sigma_{yy}^{(k)})}{\sum_{l=1}^K \pi^{(l)}\phi(y;\mu_y^{(l)},\Sigma_{yy}^{(l)})}
$$
(7)

$$
\mu_{x|y}^{(k)} = \mu_x^{(k)} + \Sigma_{xy}^{(k)} \left(\Sigma_{yy}^{(k)} \right)^{-1} \left[y - \mu_y^{(k)} \right] \tag{8}
$$

$$
\Sigma_{x|y}^{(k)} = \Sigma_{xx}^{(k)} - \Sigma_{xy}^{(k)} \left(\Sigma_{yy}^{(k)} \right)^{-1} \Sigma_{yx}^{(k)} \tag{9}
$$

A Bayesian solution to an non-linear inverse problem of solving for *x* in $y = F(x) + \varepsilon$ is given by

$$
p(x|y) \propto p(y|x)p(x). \tag{10}
$$

Instead of gradient-based solutions, we propose *Gaussian Mixture Inversion* (GMI): given $x \sim p(x)$ and $y = F(x) + \varepsilon$, the posterior density will be approximated as

$$
\widetilde{p}(x|y) = \sum_{k=1}^{K} \pi_{x|y}^{(k)} \phi(x; \mu_{x|y}^{(k)}, \Sigma_{x|y}^{(k)})
$$
\n(11)

defined as before.

Benchmark solution: Adaptive Metropolis (AM). Sample points *xt*+1 from a proposal distribution $N(x_t, C_t)$, where C_t is the covariance matrix of the chain at time *t*:

$$
C_t = cov([x_1, \ldots, x_t]), \qquad (12)
$$

accept new point with probability

$$
\alpha(x_t, x_{t+1}) = \min\left(1, \frac{\pi(x_{t+1})}{\pi(x_t)}\right) \tag{13}
$$

Further, we observe

$$
\mathbb{E}^Y \left[d_{\text{Hell}} \left(\pi(\cdot|Y), \pi_N(\cdot|Y) \right)^2 \right] \le 2 \left(d_{\text{Hell}} \left(\pi(\cdot, \cdot), \pi(\cdot, \cdot)_N \right)^2 + d_{\text{Hell}} \left(\pi_Y, \pi_{Y_N} \right)^2 \right) \tag{14}
$$

We validate our approach to posterior approximation by considering the following two synthetic example. Let

$$
x \in \mathbb{R}, \quad y = f(x) + \varepsilon \in \mathbb{R}, \quad f(x) = \sin(2x) - \cos(3x) \tag{15}
$$

Figure 8: Test case where true $x = -0.4$

2D Toy Example

Next, consider

$$
x \in \mathbb{R}^2, \quad y = f(x) + \varepsilon \in \mathbb{R}, \quad f(x) = \sin(x_1) - \cos(x_2) \tag{16}
$$

18/34

[Application to OCO-2](#page-21-0)

Example Using Simulated OCO-2 Data

- Retrieval for a single pixel. Quantity of interest: column averaged CO2 concentration, denoted XCO2.
- State vector *x* with realistic atmospheric and surface conditions.
- Simulated measurement: forward model evaluated at *x*, add synthetic measurement error.
- Radiance dimension reduction using PCA.
- Compare GMI and MCMC posteriors (with and without "model discrepancy", using a forward model emulator. See talk by Jouni Susiluoto, MS157, on Thursday!).

$$
y = F(x, b) + \varepsilon + \delta \tag{17}
$$

- *x*: operational prior, combination of priors, more comprehensive local / global distribution.
- *b*: forward model parameter uncertainty included by prescribing a distribution.
- *ε*: error model, possibility of off-diagonal elements in covariance, can be non-Gaussian.
- *δ*: model discrepancy for including model misspecification, can include new ML bias-correction, other methods for accounting "Unknown Unknowns" and spectral residuals.

To test performance with more realistic data, we perturb modeled radiances with model discrepancy adjustment:

$$
y = F_1(x, b_1) + \varepsilon + \delta, \quad \delta = F_0(x, b_0) - F_1(x, b_1). \tag{18}
$$

Left: O2-A Band radiance. Middle: a realization of measurement error *ε* \sim $\mathcal{N}(0, S_v)$. Right: a realization of model discrepancy *δ* \sim $\mathcal{N}(\mu_{\delta}, S_{\delta})$.

Example Using Simulated OCO-2 Data

(Without model discrepancy: $y_{obs} = F(x) + \varepsilon$)

Full posterior

Example Using Simulated OCO-2 Data

(With Model Discrepancy: $y_{obs} = F(x) + \varepsilon + \delta$)

Full posterior, with model discrepancy

Timing

- 10000 state vectors sampled from prior; input to forward model (emulator); added noise. Time: 104s (parallel in DGX station with 20 threads).
- Traned GMM with Julia's GaussianMixtures package: 100 iterations for learning, 20 mixture components. Time: 44s.
- Condition on measured radiance; sample 10000 realizations from conditional distribution. Time: 0.014s.
- Adaptive Metropolis ran for 1000000 iterations. Time: approx. 17000s, or 4.5h (using emulator).

- Training data: partition the Globe into clusters according to real OCO-2 measurements via self-organizing maps. For each cluster, derive marginal distribution on *x* and *δ*. Obtain $y = F(x, b) + \varepsilon + \delta.$
- Evaluate model performance against operational retrieval by left-out simulated data, and TCCON co-located real world retrievals of XCO2.

Comparison with Left-Out Data

Comparison with Real Data

Include Model Discrepancy and Prior Mean In Training

Comparison with Real Data (Revisited)

Remark: Averaging Kernels?

$$
A = (S_a^{-1} + K^T S_{\varepsilon}^{-1} K)^{-1} K^T S_{\varepsilon}^{-1} K
$$

\n
$$
= \hat{S} K^T S_{\varepsilon}^{-1} K
$$

\n
$$
= \hat{S} (K^T S_{\varepsilon}^{-1} K + S_a^{-1} - S_a^{-1})
$$

\n
$$
= \hat{S} (\hat{S}^{-1} - S_a^{-1})
$$

\n
$$
= \hat{S} \hat{S}^{-1} - \hat{S} S_a^{-1}
$$

\n
$$
= I - \hat{S} S_a^{-1}
$$
 (19)

- Wasserstein loss and cross-validation to fit GMM.
- Fitting/learning the mixture model with e.g. mixture density networks. Different basis functions, e.g. Cauchy.
- Dimension reduction for data using Autoencoders, UMAP.
- Other methods with similar capabilities: Deep Ensemble methods, VAEs, GAN-FLOW.

Thank you!

References:

1 Braverman et al. *Post hoc Uncertainty Quantification for Remote Sensing Observing Systems*, 2021

<https://epubs.siam.org/doi/10.1137/19M1304283>.

- 2 David et al: XCO2 estimates from the OCO-2 measurements using a neural network approach, 2021
	- *<https://doi.org/10.5194/amt-14-117-2021>*
- 3 Bréon et al.: On the potential of a neural-network-based approach for estimating XCO2 from OCO-2 measurements, 2022 *<https://doi.org/10.5194/amt-15-5219-2022>*
- 4 Lamminpää et al. Accelerated MCMC for Satellite-Based Measurements of Atmospheric CO2, 2019 *<https://doi.org/10.3390/rs11172061>*

This work was performed at the Jet Propulsion Laboratory, California Institute of Technology, under contract with the National Aeronautics and Space Administration. Government sponsorship acknowledged

